Exam

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Name	
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1) Radio waves	visible lic	ht, and X-rays are all types of electromagnetic radiation.	1)
Answer: T	_	False	.,
2) The frequency	of a wate	er wave gives us its height.	2)
Answer: Ti	rue 🥝	False	
3) If a new wave	arrives o	n shore every two seconds, then its frequency is 2 Hz.	3)
Answer: T	rue 🧧	False	
4) The greater th	e disturba	ance of the medium, the higher the amplitude of the wave.	4)
Answer: 0 Ti	rue	False	
5) While gravity	is always	attractive, electromagnetic forces are always repulsive.	5)
Answer: T	rue 🧧	False	
6) Changing the	electric fi	eld will have no effect on the magnetic fields of a body.	6)
Answer: T	rue 🧧	False	
7) As they move perpendicular	•	space, the vibrating electrical and magnetic fields of a light wave must move other.	7)
Answer: 0 Ti	rue	False	
8) Wave energy	can only l	pe transmitted through a material medium.	8)
Answer: T	rue 🧧	False	
	•	arough a prism, the red (longer) wavelengths bend less than the blue so forming the rainbow of colors.	9)
Answer: 0 Ti	rue	False	
0) Observations	in the X-ı	ray portion of the spectrum are routinely done from the surface of the Earth.	10)
		False	·
,		the energy is radiated uniformly in every region of the spectrum, so the black in color.	11)
Answer: T	rue 🥥	False	
2) According to	Wein's lav	v, the larger the blackbody, the shorter its peak wavelength.	12)
ŭ		False	,

False

Answer: V True

14)	According t	to Wein's lav	v, the higher the surface temperature of a star, the redder its color.	14)
	Answer:	True 💿	False	
15)	Doubling th	•	ure of a blackbody will double the total energy it radiates.	15)
	Answer:	True 🥥	False	
16)		•	increases, the frequency of peak emission also increases.	16)
	Answer: 0	True	False	
17)	The spectra emission or		th element are distinctive to that element, whether we are looking at lines.	17)
	Answer: 0	True	False	
18)	An absorpti		etrum, with dark lines crossing the rainbow of the continuum, is produced as.	18)
	Answer:	True 📀	False	
19)	An emission atomic nucl		s from an electron falling from a higher to lower energy orbital around its	19)
	Answer: 0	True	False	
20)	The shorter	a wave's wa	avelength, the greater its energy.	20)
	Answer: 0	True	False	
21)	Spectral line	es are produ	nced when an electron makes a transition from one energy state to another.	21)
	Answer: 0	True	False	
22)	In the Bohr	model of th	e atom, an electron can only exist in specific, well-defined energy levels.	22)
	Answer: 0	True	False	
23)			ydrogen atom drops from the second to the first excited energy state it sion line called hydrogen alpha.	23)
	Answer: 0	True	False	
24)	The Zeeman	n effect reve	als the presence of strong magnetic fields by the splitting of spectral lines.	24)
	Answer: 0	True	False	
25)	The broader	r the spectra	Il line, the higher the pressure of the gas that is creating it.	25)
	Answer: 0	True	False	
26)	In the Dopp	oler effect, a	redshift of spectral lines shows us the source is receding from us.	26)
	Answer: 💿	True	False	
27)	The larger t		the faster the distant galaxy is rushing toward us.	27)
	Answer:	True 🕏	False	
28)			ising in pitch, it must be approaching us.	28)
	Answer: 0	True	False	

	29) You would perceive a change in a visible light wave's amplitude as a change in its color.	29) _	
	Answer: True 🕑 False		
	30) Spectroscopy of a star can reveal its temperature, composition, and line-of-sight motion.	30)	
	Answer: True False	_	
	31) The Doppler effect can reveal the rotation speed of a star by the splitting of the spectral lines.	31)	
	Answer: True 🗸 False		
MUL	TIPLE CHOICE. Choose the one alternative that best completes the statement or answers the question.		
	32) Which of these is NOT a form of electromagnetic radiation?	32)	
	A) X-rays in the doctor's office		
	B) DC current from your car battery		
	C) ultraviolet causing a suntan		
	D) light from your camp fire E) radio signals		
	Answer: B		
	Albwei. D		
	33) A wave's velocity is the product of the	33)	
	A) frequency times the wavelength of the wave.	_	
	B) frequency times the period of the wave.		
	C) period times the energy of the wave.		
	D) amplitude times the frequency of the wave.		
	E) amplitude times the wavelength of the wave.		
	Answer: A		
	34) Consider this diagram. Which statement is true?	34) _	
	$ \begin{array}{c} & \downarrow \\ $		
	A) The amplitude is 8 and the wavelength is 6.B) The amplitude is 8 and the wavelength is 12.C) The amplitude is 6 and the wavelength is 4.		
	D) The amplitude is 4 and the wavelength is 12.		
	E) The amplitude is 4 and the wavelength is 6.		
	Answer: E		
	35) If a wave's frequency doubles and its speed stays constant, its wavelength	35)	
	A) is now 4× longer.		
	B) is halved.		
	C) is also doubled.		
	D) becomes 16× longer.E) is unchanged, as c is constant.		
	Answer: B		

36) The speed of light in a vacuum is	36)	
A) 186,000 miles per hour.		
B) $h = E/c$.		
C) 300,000 km/sec.		
D) 768 km/hour.		
E) none of the above.		
Answer: C		
37) Which of these is the same for all forms of electromagnetic (E-M) radiation in a vacuum?	37)	
A) frequency		
B) wavelength		
C) speed		
D) amplitude		
E) photon energy		
Answer: C		
38) The two forms of electromagnetic (E-M) radiation that experience the least atmospheric opacity are	38)	
A) microwaves and radio waves.		
B) X and gamma radiation.		
C) ultraviolet and infrared waves.		
D) visible light and infrared waves.		
E) visible light and radio waves.		
Answer: E		
39) The radiation our eyes are most sensitive to is the color	39)	
A) black at 227 nm.		
B) violet at 7,000 Angstroms.		
C) red at 6563 Angstroms.		
D) yellow-green at about 550 nm.		
E) blue at 4,321 nm.		
Answer: D		
40) Medium A blocks more of a certain wavelength of radiation than medium B. Medium A has a	40)	
higher		
A) opacity.		
B) transparency.		
C) clarity. D) seeing.		
E) albedo.		
Answer: A		
Allswei. A		
41) In the Kelvin scale, absolute zero lies at	41)	
A) -373 degrees C.		
B) 273 degrees C.		
C) zero K.		
D) Both A and B are correct.E) Both A and C are correct.		
E/ DOTH A AND CAIS CONSCI.		

Answer: C

42) What is true of	a blackbody?				42)	
A) Its energy	is not a continuum.				_	
	mplete absence of the					
	peaks at the waveler					
-		•		oubled in wavelength.		
E) It appears	black to us, regardle	ss of its temperatu	re.			
Answer: C						
43) What is the nam	ne of the temperature	e scale that places z	zero at the point whe	re all atomic and	43)	
molecular motion	on ceases?					
A) Ransom						
B) Fahrenhei	t					
C) Kelvin						
D) Centigrad	е					
E) Celsius						
Answer: C						
	y radiated by a blackl				44)	
	root of its temperatu	ire.				
	of its temperature.					
•	of its temperature.					
·	power of its tempera					
•	e root of its temperatu	ure.				
Answer: D						
_	emperature of a blacl		of 3 will increase its e		45)	
A) 3.	B) 6.	C) 9.	D) 12.	E) 81.		
Answer: E						
46) If a star was the	same size as our Sur	n, but was 81 times	s more luminous, it m	nust be	46)	
A) twice as h	ot as our Sun.				_	
B) three time	s hotter than the Sun	l.				
C) four times	hotter than the Sun.					
·	s hotter than the Sun.					
E) 81 times h	otter than the Sun.					
Answer: B						
47) The Sun's obser	ved spectrum is				47)	
A) a continuu	um with absorption I	ines.			_	
. •	rption lines on a blac	•				
	um with no lines, as s	-	00W.			
,	um with emission lin					
E) only emis	sion lines on a black l	background.				

Answer: A

48) The element first found in the Sun's spectrum, then on Earth 30 years later, is	48)
A) technicum.	
B) hydrogen.	
C) helium.	
D) aluminum.	
E) solarium.	
Answer: C	
7 (1500).	
49) A jar filled with gas is placed directly in front of a second jar filled with gas. Using a spectroscope	e 49)
to look at one jar through the other you observe dark spectral lines. The jar closest to you contain:	s
A) gas at the same temperature as the other jar.	
B) the cooler gas.	
C) gas at very high pressure.	
D) the hotter gas.	
E) the exact same gas as the other jar.	
Answer: B	
50) Which of these is emitted when an electron falls from a higher to lower orbital?	50)
A) a neutrino	
B) a positron	
C) a graviton	
D) a photon	
E) another electron	
Answer: D	
51) In Bohr's model of the atom, electrons	51)
A) can be halfway between orbits.	JI)
B) are not confined to specific orbits.	
·	
C) are spread uniformly through a large, positive mass.	
D) move from one orbit to the next orbit in many small steps.	
E) only make transitions between orbits of specific energies.	
Answer: E	
52) In general, the spectral lines of molecules are	52)
A) less complex than those of atoms.	
B) only absorption lines.	
C) more complex than those of atoms.	
D) the same as the atoms they contain.	
E) nonexistent.	
Answer: C	
53) Electromagnetic radiation	53)
A) has only the properties of waves.	
B) can behave both as a wave and as a particle.	
C) can only travel in a dense medium.	
D) is the same as a sound wave.	
E) has nothing in common with radio waves.	
Answer: B	

54) In a hydrogen atom, a transition from the 2nd to the 1st excited state will produce	54)
A) the bright red Balmer alpha emission line.	
B) three different emission lines.	
C) a dark absorption line.	
D) an ultraviolet spectral line.	
E) no emission line.	
Answer: A	
55) For hydrogen, the transition from the first to third excited state produces	55)
A) a blue green absorption line.	, <u> </u>
B) an ultraviolet line.	
C) a red emission line.	
D) an infrared line.	
E) a violet emission line.	
·	
Answer: A	
56) The observed spectral lines of a star are all shifted towards the red end of the spectrum. Which	56)
statement is true?	,
A) The star has a radial velocity towards us.	
B) The second law of Kirchhoff explains this.	
C) The star is not rotating.	
D) This is an example of the photoelectric effect.	
E) This is an example of the Doppler effect.	
Answer: E	
Allswei. E	
57) If a source of light is approaching us at 3,000 km/sec, then all its waves are	57)
A) blueshifted by 1%.	
B) blueshifted out of the visible spectrum into the ultraviolet.	
C) redshifted by 1%.	
D) redshifted out of the visible into the infrared.	
E) not affected, as c is constant regardless of the direction of motion.	
Answer: A	
Allower. A	
58) If the rest wavelength of a certain line is 600 nm, but we observe it at 594 nm, then	58)
A) the source is getting 1% hotter as we watch.	,
B) the source is approaching us at 0.1% of the speed of light.	
C) the source is receding from us at 10% of the speed of light.	
D) the source is approaching us at 1% of the speed of light.	
E) the source is spinning very rapidly, at 1% of the speed of light.	
Answer: D	
Allswei. D	
59) According to the Zeeman effect, the splitting of a sunspot's spectral lines is due to	59)
A) temperature variations.	, <u> </u>
B) a Doppler shift.	
C) their radial velocity.	
D) their rapid rotation.	
E) their magnetic fields.	
Answer: E	

60)	second; this is a typic	cal carrier frequency	for FM (frequency	· ·	·	60)
	A) millihertz Answer: D	B) gigahertz	C) hertz	D) megahertz	E) kilohertz	
61)	According to Wein's	law, the wavelength	n of the peak energ	y will be if th	e temperature of	61)
	the blackbody is dou A) doubled B) quadrupled C) quartered D) unchanged E) halved	bled.				
	Answer: E					
62)	The Sun's blackbody A) ultraviolet Answer: E	curve peaks in the _ B) X-ray	portion o C) radio	f the spectrum. D) infrared	E) visible	62)
63)	The common elemen A) hydrogen. Answer: A	t with bright red, blu B) nitrogen.	ue-green, and viol C) oxygen.	et emission lines is D) carbon.	E) helium.	63)
64)	 The most energetic p A) infrared. B) radio. C) X-rays. D) gamma rays. E) visible. Answer: D 	notons are				64)
OLIODE A						
		·	•	ch statement or answe	·	
65)	The distance from a N Answer: amplitude	wave's crest to its un	idisturbed position	is the	65) _	
66)	The product of the w	avelength times the	frequency of a wa	ve is its	66)	
	Answer: velocity					
67)	A wave with a period Answer: 100	d of .01 seconds has	a frequency of	Hz.	67) _	
68)	A wave with a freque Answer: one-half se	_	ve a period of		68) _	
69)	An FM station broad	casts at a frequency	of 100 MHz. The v	vavelength of its carrier	wave is 69)	
	Answer: 3 meters					
70)	In electromagnetic w Answer: perpendicu		d magnetic fields v	/ibrate to each	other. 70) _	

71)	A featureless spectrum, such as a rainbow, is said to be	71)	
	Answer: continuous		
72)	Stars that appear blue or white in color are than our yellow Sun.	72)	
	Answer: hotter		
73)	Knowing the peak emission wavelength of a blackbody allows you to determine its	73)	
	Answer: temperature		
-	Stefan's law notes that total energy radiated is proportional to the power of the temperature of the blackbody.	74)	
	Answer: fourth		
	A dense, hot body will give off a(n) spectrum.	75)	
	Answer: continuous		
	Fraunhofer was the German astronomer who first noted lines in the Sun's spectrum.	76)	
	Answer: absorption		
77)	The common element discovered in the Sun's spectrum before it was found here is	77)	
	Answer: helium		
78)	When an electron moves from a lower to a higher energy state, a photon is	78)	
	Answer: absorbed		
79)	An electron has a electric charge.	79)	
	Answer: negative		
80)	The energy of the photon depends on its	80)	
	Answer: frequency or wavelength.		
-	Why can't we be certain that the Andromeda Galaxy exists today?	81)	
	Answer: Since it lies 2.5 million light-years distant, the most recent image we have is still 2.5 million years out of date, so we cannot prove it is still there. It probably is, though.		
82)	How do sound and light waves differ?	82)	
	Answer: Sound waves travel much slower, and need a physical medium, such as air, to be transmitted. Light travels best in the vacuum of space.		
-	An AM station is broadcasting at 980 kHz, while an FM station up the road is assigned 98 MHz. How do their carrier waves compare?	83)	
	Answer: As the frequency of the FM station is 100 times higher than the AM station, the FM carrier wave must be 100 shorter in wavelength.		

	No one can hear you scream (or fire a weapon) in space, regardless of the Hollywood	84)	
	special effects. Explain why.		
	Answer: Sound waves must travel though a material medium, and cannot pass through a vacuum. The blast might be seen, but the boom will not be heard.		
	What two regions of the electromagnetic spectrum are best utilized by ground-based astronomers, and why?	85)	
	Answer: The atmosphere is opaque to most radiation except visible and radio waves.		
	How can you determine the distance to a spacecraft from the time it takes its radio signal to reach Earth?	86)	
	Answer: In a vacuum, all electromagnetic radiation, including radio waves, travel at the sam speed: 300,000 km/s. Measuring the time it takes the radio signal to reach us and multiplying by 300,000 km/s gives the distance to the spacecraft.	ie	
	Newton found that when light passed through a prism, it was dispersed into the component colors. Which bent the least, and why?	87)	
	Answer: The red waves are bent less by the glass than are the other colors because they have the longest wavelength. Shorter wavelengths bend more than longer wavelengths.		
88)	What do infrared and ultraviolet waves have in common? How do they differ?	88)	
	Answer: Both are forms of electromagnetic radiation, both travel at c in a vacuum, and both are largely absorbed by our atmosphere. They differ greatly in frequency, wavelength, and photon energy, however, with UV much more energetic than IR.		
89)	What do gamma rays, X-rays, light, and radio waves all have in common?	89)	
	Answer: While they vary widely in wavelengths and frequencies, they are all forms of electromagnetic radiation and all travel at c , the speed of light, in a vacuum.		
90)	How does human vision's peak in color sensitivity relate to the Sun?	90)	
	Answer: Our eyes are tuned to utilize best the type of radiation our star produces the most o and yellow lies in the middle of the visible spectrum.	f,	
91)	Give at least two advantages of the Kelvin temperature scale for astronomers.	91)	
	Answer: It is an absolute scale, so there are never any negative readings. Wein's and Stefan's laws are only mathematically correct if Kelvin temperatures are used.		
	The Great Nebula in Orion, M-42, is a low-density cloud of hot gas. Use Kirchhoff's laws to describe its spectrum.	92)	
	Answer: Kirchhoff's second law notes that a hot thin gas will create an emission spectrum of bright lines through the spectroscope.		
	According to Kirchhoff's first law why do dense, hot bodies create the type of spectrum they do?	93)	
	Answer: Kirchhoff's first law states that a dense, hot medium emits light of all wavelengths, creating a continuous spectrum.		
	If the magnetic fields are very strong, such as around sunspots, how are spectral lines affected by the Zeeman effect?	94)	

Answer: A strong magnetic field will cause the lines to appear split apart.

95)	State the re	lationship between frequency, photon energy, and wavelength.	95)
		he higher the frequency, the greater the energy the photon carries, but the shorter s wavelength.	
96)	Explain hov	w the Zeeman effect allows us to study stellar magnetic fields.	96)
	m	he Zeeman effect causes spectral lines to appear split into two. This tells us nagnetic fields are present. The greater the observed splitting, the stronger the nagnetic fields are.	
97)	Explain hov	w Bohr's model creates emission and absorption lines in the spectrum.	97)
	tr Ea	ohr's model has the electron orbitals quantized into discrete energies. Each upward ransition to a higher energy state produces an absorption line (energy is absorbed). ach downward transition produces an emission line (energy is emitted). The nergy absorbed or emitted is exactly equal to the difference in energy levels.	
98)	What infor	mation about a star can be inferred from its Doppler shift?	98)
	Answer: TI	he Doppler shift gives the star's radial velocity, either towards or away from us.	
99)	-	ar system is one with two stars orbiting each other. How can the Doppler effect find binary stars whose orbital plane is along our line of sight and determine ds?	99)
	of	ss the two stars orbit each other rapidly, one will approach us, creating a blueshift fits spectral lines, while its retreating companion shows a redshift. The time to go brough two splits and recombinations of their lines is their orbital period.	
100)	Explain wh	nat types of information can be obtained from a line spectrum.	100)
	sp	he element which created it, the line-of-sight velocity of the source, its rotation beed, temperature, the pressure of the gas emitting the radiation, and even its hagnetic field may also be found.	
101)	If we increa	ased the pressure in a gas, how will its spectral lines be affected?	101)
	Answer: Tl	he lines will broaden (or even disappear if the density becomes too great)	
102)	Contrast th	e speeds of sound and light in watching a flash of lightning, then listening for r to follow.	102)
	fe	ight travels at 300,000 km/sec, so the flash of light is almost instantaneous from a ew miles away; sound travels at about a fifth of a mile per second, so if the thunder bllows the lightning by five seconds, the bolt hit about a mile away.	
103)	How can W	Vein's law be used to determine the temperature of a star?	103)
		areful analysis of the blackbody curve of the star's entire radiation spectrum will eveal a peak that is unique to a given temperature. Basically, the bluer the star's	

radiation, the hotter its surface will be.

	3	uld a hotter star appear blue-white while a cooler star appear red or not be visible	104)
	at all?		-
	Answer:	Stefan's law notes that the higher the temperature, the more luminous the body is, so such stars produce great amounts of visible light. The hotter the star the shorter the wavelength it peaks at. A star that emits light across the entire visible spectrum would appear white. One that peaked beyond the visible would appear blue-white. A cooler star may peak in the red part of the spectrum, or even in the infrared.	
		s Stefan's law and a knowledge of Earth's history tell us that the Sun's ure cannot have varied much in the last 3.5 billion years?	105)
	Answer:	Since even a small change in temperature, raised to the fourth power, would result in a large change in the total solar energy radiated, if the Sun had cooled much, our oceans would have frozen and life would have ceased to exist here.	
106)	Explain t	he appearance of the Sun's spectrum, as noted by Fraunhofer.	106)
	Answer:	The Sun is dense, and gives rise to a continuous spectrum, peaked in the color yellow as dictated by the 5800K temperature of its surface. Then the cooler, less dense gas above the surface absorbs some of the energy in transit, revealing its composition by the particular absorption lines we observe from Earth.	
107) How does the energy of a water wave differ from the energy of a photon?			107)
	Answer:	Amplitudes of sound (and water) waves can differ greatly and still have the same wavelength and frequency, as they are the result of the motions of large numbers of molecules. For photons, the energy is quantized, so that each photon of a given wavelength must carry the same amount of energy.	
	_	we know that the red Balmer emission line in hydrogen represents a smaller leap than the violet line?	108)
	Answer:	Red light has a longer wavelength than violet light; therefore a red photon contains less energy than a violet one. Since the photon given off when an electron's energy level changes has an energy equal to the energy difference between the two levels, the less energetic photon represents a smaller difference.	
109)	Give an e	example of the Doppler effect being used in a baseball game.	109)
	Answer:	The Doppler "gun" can focus on the motion of the baseball, and give us the speed that the pitcher is delivering it to the plate.	
110)	Give and	explain an example of the use of the Doppler effect on the highway.	110)
	Answer:	The radar gun of a highway patrolman sends out a pulsed beam to be reflected back, thus giving the speed of your car and perhaps netting you a ticket.	
111)	How can	the Doppler effect be used to determine if a storm is forming into a tornado?	111)
	Answer:	Radar can determine the distance to a storm cloud. Since a tornado rotates very rapidly. Doppler radar can measure the difference in velocity between the two sides	

of the storm to determine if it is rotating.

112) Explain how the Doppler effect has been used to detect invisible planets orbiting other Sun-like stars.

112) _____

Answer: The planets are massive enough to pull their star slightly off course as they orbit from one side to the other, producing a cycle of redshifts and blueshifts that allow us to deduce that the planet is present, and how long it takes to orbit its star.