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### **Chapter 02: Knowing the Heavens**

- 1. Which of the following was not obtained by people of ancient civilizations from observations of the night sky?
  - a. the relative distances of the Sun, the Moon, and stars from Earth
  - b. timing information, both daily and yearly
  - c. patterns of stars in the sky about which stories and myths were devised
  - d. directions for navigation

ANSWER: a

- 2. Two opposite sides of the Egyptian pyramids are aligned with which direction(s) on the horizon?
  - a. northeast to southwest
  - b. midsummer sunrise
  - c. north-south
  - d. rising of the star Sirius

ANSWER:

- 3. Meaningful observations of the sky by ancient peoples:
  - a. were, of necessity, restricted to the western hemisphere.
  - b. were, of necessity, restricted to the regions north of the equator.
  - c. were, of necessity, restricted to the regions south of the equator.
  - d. were made in regions throughout the world.

ANSWER:

- 4. Which ancient culture discovered a 584-day pattern in the motions of Venus?
  - a. Native Americans in the southwestern United States
  - b. Mayan civilization in Central America
  - c. Stonehenge builders in England
  - d. Observers in ancient Egypt

ANSWER: b

- 5. In modern astronomy, the constellations are:
  - a. clusters of stars that are held together by the mutual gravitational attractions of the individual stars in the cluster.
  - b. nearby galaxies to which astronomers have given specific names.
  - c. 12 regions of sky through which the Sun, the Moon, and planets move as seen from Earth.
  - d. 88 regions of sky, covering the entire sky.

ANSWER:

- 6. Constellations are generally made up of stars that are:
  - a. moving through space together.
  - b. about the same age.

c

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c. separated by relatively small distances (compared to their distances from us).	
d. merely located along more or less the same line of sight as viewed from Earth.	
ANSWER:	d
7. The constellations:	
a. are 88 in number and cover the entire sky.	
b. that the ancients imagined are constantly being augmented by newly invented costars are being discovered.	onstellations as new
c. are of historical interest only and play no role at all in modern astronomy.	
d. consist of groups of stars that are all about the same distance from us.	
ANSWER:	a
8. Describing a star as being in the constellation Cygnus (the Swan) tells a modern astronal in a distant galaxy located in a particular direction from Earth.	onomer that the star is:
b. inside our solar system.	
c. somewhere in a particular region of sky having definite boundaries.	
d. one of a set of bright stars that make up a particular "picture" in the sky.	
ANSWER:	c
9. If a star is described as being in the constellation Leo, a modern astronomer knows that a. somewhere in a particular region of sky having definite boundaries.	nat it is:
b. in a distant galaxy located in a particular direction from Earth.	
c. inside a region of the sky bounded by two lines of right ascension in the sky.	
d. one of a few individual bright stars that make a picture (of a lion) in the sky.	
ANSWER:	a
10. Which of the following statements correctly describes the relationship between star a. Only those stars that were visible to the ancient Greeks are located in constellat b. Only the brighter stars are in constellations.	
c. Only stars within the zodiac close to the ecliptic, Earth's orbital plane, are located	ed in constellations.

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d. Every star is located in a constellation.

ANSWER:

- 11. If the unaided human eye is sensitive enough to see about 6000 of the stars in the entire sky, about how many stars would be seen at one time on a given night from a single location from which the horizon is completely visible around the observer?
  - a. about 3000

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- b. only 1000 because the rest are hidden by Earth
- c. depends on the observer's latitude; observers at the poles will see 6000, while equatorial observers will see only one-half of this number, or 3000
- d. about 6000

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ANSWER:				a
<ul><li>a. If you</li><li>b. Only</li><li>c. Astro</li></ul>	n point random if you point in nomers are see	ing statements about construction in the state vicinity of one of the beking to discover new constructions and the stantly inventing new constructions.	sky, you are pointing righter stars are you ellations.	
<ul><li>a. Lec</li><li>b. Urs</li><li>c. Urs</li></ul>	o, the lion, consa Major, the Csa Minor, the I	taining the bright star Regu	lus. right star Polaris.	al pole in the northern hemisphere is:
14. The Winte	er Triangle, a g a. b. c. d.	group of three bright stars in Betelgeuse. Vega. Castor. Polaris.	i the winter sky, co	onsists of Sirius, Procyon, and:
15. The Sumr	ner Triangle, a a. b. c. d.	group of three bright stars Betelgeuse. Vega. Pollux. Polaris.	in the summer sky	y, consists of Deneb, Altair, and:
16. Diurnal ma. b. c. d. ANSWER:	precession of revolution of the	ts in the sky is caused by the fearth's axis. If Earth around the Sun. The Moon across the sky. The Earth on its axis.	e:	d
_	•	ojects across the sky is caus a around the Sun.	ed by the:	

b. rotation of the whole celestial sphere of stars around the fixed Earth.

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c. rotation of Ear	th on its axis.		
d. motion of the s	solar system around the ga	alaxy.	
ANSWER:			c
18. A star rises at 8 P.M	I., moves across the sky (c	crossing high overhead	), and then sets at:
a			
b	o. 2 A.M.		
c	e. 8 A.M.		
d	l. noon.		
ANSWER:			c
instant your friend on the a. is good b. is good b.	from the east coast of the Une west coast will find that just rising.  just setting.  se a few hours earlier.		rises above the horizon. At this same
d. ha	s not yet risen.		
ANSWER:			d
<ul><li>a. motion of the</li><li>b. rotation of Ea</li><li>c. relative motion</li></ul>	served east-to-west motion. Moon and planets across arth on its axis.  Ons of stars with respect to Earth around the Sun.	the sky.	
ANSWER:			b
sky over 1 night. This is	s caused by the:	osition on Earth gradua	ally shifts from east to west across the
1	c motions and winds.	.1 .	
	Earth about its own north-		
	he Moon and planets acro	oss the sky.	
	Earth around the Sun.		1.
ANSWER:			b
•	erved motion in the night f Earth around the Sun.	sky is produced by the	x:
b. motion of sta	ars with respect to each of	ther in the sky.	

ANSWER:

rotation of Earth on its axis.

motion of the planets along their orbits around the Sun.

d

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<b>C</b> 1	to the: constellations from east to west from isible at 11 P.M. in May than at 11	
b. apparent motion of the Sun alor	ng the ecliptic over the course of a	year.
d. gradual motion of the constellar	in the sky as it runs through its phations from east to west across the skisible at 4 A.M. than at 10 P.M. on	ky each night, resulting in
ANSWER:		d
24. Over the course of 1 night, an obsergradually move from east to west across a. inherent rotation of the	ss the sky. This is caused primarily	
<ul><li>b. wind.</li><li>c. motion of Earth around</li></ul>	the Sun	
c. motion of Earth around d. rotation of Earth around		
ANSWER:	a its own dats.	d
25. What basic pattern do stars seem to celestial pole for a period of several ho		or photograph) stars near the north
a. circles, with the north celesti	al pole at the center	
b. spirals, as the stars move wh	ile Earth rotates	
c. almost straight lines, rising fi	rom the horizon towards the zenith	
d. ellipses, with the north pole a	at one focus	
ANSWER:		a
26. When we watch the nighttime sky,	we find that:	
<ul> <li>a. the stars and constellations rem because they are so far away.</li> </ul>	ain fixed in our sky, not rising or se	etting in a time as short as 1 night
b. most stars and constellations slo	owly rise in the west, pass overhead	d, and set in the east.
c. all stars and constellations reach	h their highest point in the sky at m	idnight.
d. most stars and constellations slo	owly rise in the east, pass overhead	, and set in the west.
ANSWER:		d
27. Which way are you moving with re	espect to the stars during the rotation	n of Earth?
a. east	tward	
b. sou	thward	
c. wes	stward	
d. nor	thward	
ANSWER:		a

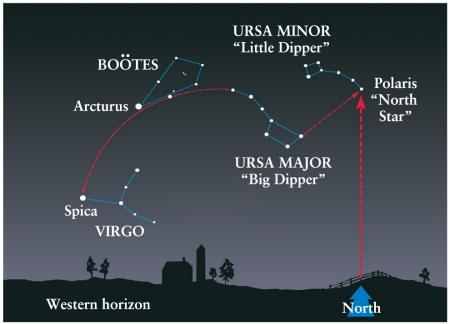
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28. With respect to the stars, the rotation of Earth carries you toward the:

- a. south.
- b. east.
- c. west.
- d. north.

ANSWER: b

29. If the stars Polaris and Arcturus (as shown in Figure 2-6 of *Universe*, 11th ed.), are known to be 71° apart, how far away from Polaris is the closest star in Ursa Major?

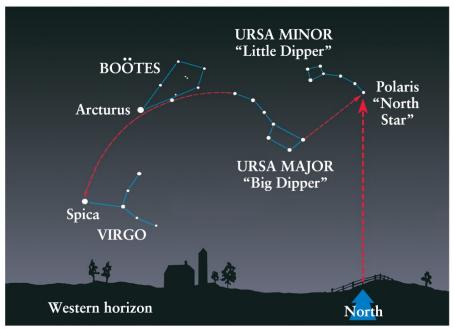


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- a. about 25°
- b. about  $2.5^{\circ}$
- c. about 250°
- d. about 7.1°

ANSWER:

30. If the stars Polaris and Spica are 101° apart (as shown in Figure 2-6 of *Universe*, 11th ed.), how far away from Polaris is the closest star in Ursa Major, from our view?



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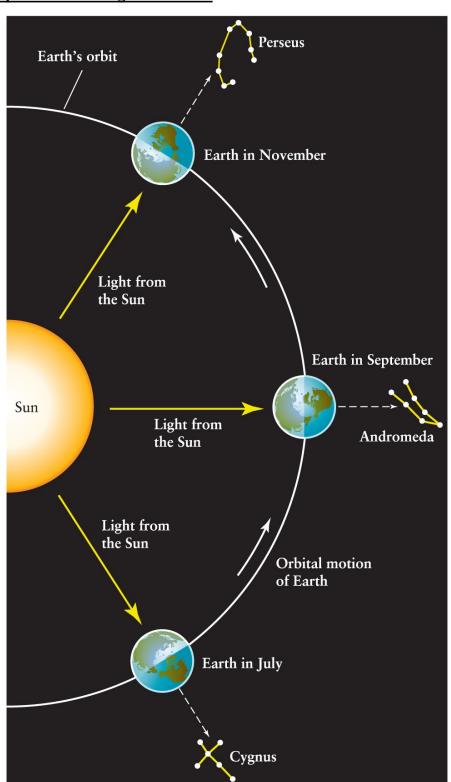
- a. about  $2.5^{\circ}$
- b. about 260°
- c. about 77°
- d. about 25°

ANSWER:

- 31. In late September, Andromeda appears high in the sky at midnight. Six months later Virgo appears high in the sky at midnight. Where is Andromeda at this time?
  - a. It is still high in the sky at midnight.
  - b. It has moved to the western horizon.
  - c. It has moved to the eastern horizon.
  - d. It is high in the sky at *noon* and is thus not visible.

ANSWER:

32. Perseus appears high in the sky at midnight in November. Andromeda appears high in the sky at midnight in September. (See Figure 2-5 of *Universe*, 11th ed.) Where is Andromeda at midnight in November?



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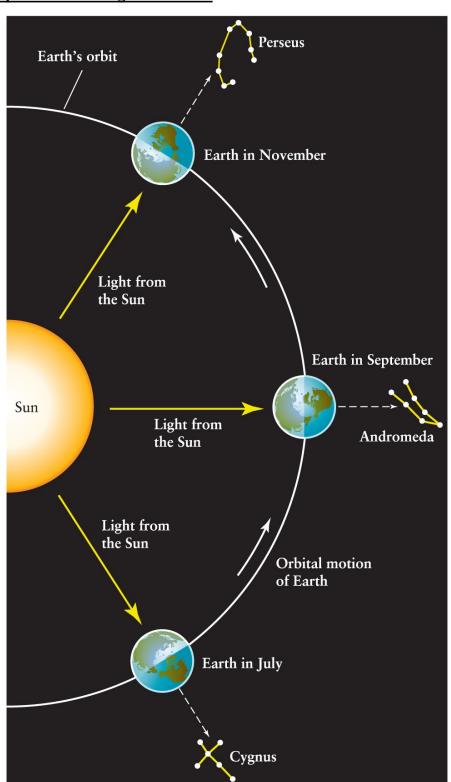
- a. It remains high in the center of the sky.
- b. It will be near the eastern horizon.

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- c. It will be near the western horizon.
- d. It will be high in the sky at noon and thus will not be visible.

ANSWER:

33. Cygnus appears high in the sky at midnight in July. Andromeda appears high in the sky at midnight in September. (See Figure 2-5 of *Universe*, 11th ed.) Where is Andromeda at midnight in July?



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- a. It is still high in the sky at midnight.
- b. It has moved to the western horizon.

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c.	It has moved to	the eastern horizon.		
d.	It is high in the	sky at noon and is th	us not visible.	
ANSWER:				c
34. Sirius	s the brightest sta	ar in the night sky. Ye	ou watch it set behind the	western horizon about 2 A.M. in the
middle of	February. About		n the middle of March?	
	a.	10 P.M.		
	b.	midnight		
	c.	2 A.M.		
	d.	4 A.M.		
<i>ANSWER:</i>				b
35 People	in which one of	the following cities e	xperience sunrise first?	
os. i copic	a.	New York	Aperience sum ise mist.	
	b.	San Francisco		
	c.	Chicago		
	d.	Denver		
ANSWER:				a
36. Which	_		lways remain fixed in plac	e relative to an observer's horizon?
a.	summer solstice			
b.	north celestial p			
c.	-	e celestial equator con	itacts the horizon	
d.	zenith			
ANSWER:				a
37. Which	of the following	points remains fixed	in the sky relative to an ol	oserver's horizon?
a.	north celestial p		in the ship relative to the ex	
b.	_	stant star (e.g., Betel	geuse, in Orion)	
c.	vernal equinox		,	
d.	winter solstice			
ANSWER:				a
20 3371 : 1	6.4 6.11	1:	1. 4 1 1 1	1 11 2
	_	directions remains in	xed in the sky relative to a	n observer's horizon?
a. 5	zenith	a Maan at maan assa	r 1 manth	
b.		e Moon at noon, over		
c.		e Sun at noon, over 1	year	
d.	autumnal equi	шох		
<i>ANSWER:</i>				a

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	a.	vernal equinox.	
	b.	zodiac.	
	c.	celestial equator.	
	d.	zenith.	
ANSWER:			d
your zenith	at some t	ling on the equator, which one of the following positions in the sky wime in one 24-hour period? of the Sun at summer solstice	ould pass through
	_	of the north celestial pole, or perpendicular to the celestial equator	
	=	of the ecliptic pole, or perpendicular to the ecliptic plane	
-		of the vernal equinox, or 0 hours right ascension, 0° declination	
ANSWER:	position	or the verial equition, or a neuro right assertion, a destination	d
a. all		e overall sky is north of the celestial equator, that is, in the northern he definition -half	emisphere?
c. less	s than on	e-half, because of the tilt of the equator to the ecliptic plane	
d. mo	re than o	one-half, because of the precession of the poles	
ANSWER:			b
-		ts appearance to a person in mid-latitude northern latitudes, how will bserver in Australia?	the constellation
a. exac	tly the sa	ame since the stars of Orion are very far away from Earth	
b. upsi	de down	but with the same orientation of stars	
c. upsi	de down	and inverted left to right, a mirror image of that seen in the northern l	hemisphere
d. same	e way up	but inverted left to right	
ANSWER:			b
43. The cele	stial equ	ator is defined as the:	
		to space of Earth's equator.	
b. line	traced in	n our sky by the Moon each month against the background stars.	
		n our sky by the Sun over 1 year against the background stars.	
	d of cons	stellations through which the Sun and Moon move in our sky.	
ANSWER:			a
44. Which o		owing lines or points is always directly over your head, no matter wh	ere on Earth you go?
	a.	zenith	
	b.	celestial equator	
	c.	ecliptic	
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		90° north declinat	tion	
ANSWER:				a
45. If you poi same directio		•	and point there again 45 days lat	ter, you will have pointed twice in the
	a.	Moon.		
	b.	Sun.		
	c.	fixed s	stars.	
	d.	horizor	n.	
ANSWER:				d
year will mai a. a fixe b. an an c. parall	vertically vertically toward th vertically estanding of ntain which ad angle of gle to Earth	y above a point on above an observe an at noon. y above the North on Earth's equator h of the following 23.5° to the spin h's spin axis that was spin axis	r, your zenith (the vertical direct g alignments? axis of Earth will vary between 0° and 23.5°	b tion above your head) over a period of to over a period of 6 months
d. perpe	endicular to	Earth's spin axis	5	d
a. can be b. is an e of Ear merid c. rotate	e used to as extension or th's equator ian through s along wit	ssign coordinates of the latitude-long or, and the location the Greenwich, Eng	on for the zero of right ascension gland.	ne celestial equator is an extension is an extension of the prime
ANSWER:				a

49. If you view the sky from 20° north of the equator, the north circumpolar stars (which do not rise and set) would be:

- a. the stars within  $20^{\circ}$  of the horizon.
- b. the stars within 20° of Polaris, the North Star.
- c. the stars within 20° of the zenith (the point directly overhead).
- d. nonexistent, since all the stars rise and set when viewed from 20° north of the equator.

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ANSWER:			b
<ul><li>a. a path betwee</li><li>b. the celestial e</li><li>c. a variable path</li></ul>	n across the sky within the zodiac	e observer's celestial mer	
d. the ecliptic place of the description of the des	ane.		b
t passes through the:	sky will reach its highest point to	o a particular observer in	the northern hemisphere when
b. zo	lestial equator. diac.		
	liptic plane. gion of the sky due south.		
ANSWER:			d
<ul><li>b. exactly per</li><li>c. above Eart</li></ul>	star," is at present:  f the north celestial pole.  pendicular to the ecliptic plane (eth)  h's magnetic pole.  t the north celestial pole.	ecliptic pole).	
ANSWER:	, the north colestial pole.		a
<ul><li>a. a slow drift</li><li>b. a wobble ba</li><li>c. a small circ</li><li>d. zero; there</li></ul>	he apparent motion of the pole state across the sky.  ack and forth in a straight line.  le with a radius of less than 1° in its no motion of the pole star by definition.	about 24 hours.	`a day is:
ANSWER:			c
a. time of your b. time of do c. observer's	ear.	լuator).	estial pole and an observer's
ANSWER:	-	•	c

55. To navigators in the northern hemisphere, their latitude in degrees is equal to:

a. 15 times the number of hours since the sun set.

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b. th	e angle betwe	een their meridian and the north celestial pole.	
	_	een the north celestial pole and their zenith.	
d. th	ie angle betwe	een the north celestial pole and their northern horizon.	
ANSWER:		-	d
56. The an	ole between a	n observer's horizon and the north celestial pole is governe	d by:
701 2110 1111	a.	latitude.	<i></i>
	<b>b</b> .	local time.	
	c.	sidereal time.	
	d.	longitude.	
ANSWER:			a
- a m 1	1 1		4 1 2 1 1 2 1 1 2
	_	between the northern horizon of a fixed observer and the no	orth celestial pole is equal to:
a. L	the observer	r's longitude.	
b.			
C.	_	cension of the vernal equinox.	
d.	a variable v	alue, depending on the time of year.	
<i>ANSWER:</i>			ь
		star straight overhead has a right ascension of 2 hours and a right ascension of:	30 minutes. One hour later
	a. 1 h	our and 30 minutes.	
	b. 2 h	ours and 30 minutes.	
	c. 3 h	ours and 30 minutes.	
	d. 4 h	ours and 30 minutes.	
ANSWER:			b
	e year) for an	leclination will stars be circumpolar (i.e., be visible at any tobserver at 30° north latitude?	time of the night on any
	a. 60° to	90°N and 60° to 90°S	
	b. 60° to	90°N	
	c. 30° to	90°N	
	d. $0^{\circ}$ to	30°N	
ANSWER:			ь

60. Over what range of declination will stars be circumpolar (i.e., be visible at any time of the night on any night of the year) for an observer at the South Pole?

- a.  $0^{\circ}$  to  $90^{\circ}$ N
- b. only stars at the south celestial pole
- c. 90°S to 90°N (the whole sky)

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d.	0° to 90°S	
ANSWER:		d
	ngles used by astronomers to define the positable anywhere on Earth are:	tion of a star in the sky and define a coordinate
a.	azimuth and elevation.	
b.	latitude and longitude.	
c.	horizontal and vertical angles.	
d.	right ascension and declination.	
ANSWER:		d
62. The celest a.	ial coordinates that together describe a star's right ascension and declination.	position precisely and unambiguously are:
b.	right ascension and sidereal time.	
c.	sidereal time and latitude.	
d.	longitude and latitude.	
ANSWER:		a
	nation angle between the north celestial pole	and the celestial equator is:
	60°.	
	9° because of the displacement of the Pole St	tar.
	80°.	
	0°.	ı
ANSWER:		d
a. positio	nation of a star is a measure of its: on above the observer's horizon along a great of rising above the horizon.	t circle passing through the observer's zenith.
	on along the celestial equator.	
d. positio		ng a great circle passing through the north and
ANSWER:		d
65. The declin	nation of a star in the sky is defined as the ang	gle between the:
a. Sun ar	nd the star, measured along the ecliptic.	
	al equator and the star, measured along a gre of the galaxy and the star, measured along the	eat circle passing through both celestial poles. the galactic equator.
d. vernal	equinox and the star, measured along the ce	elestial equator.
ANSWER:		b

66. The right ascension of a star is one coordinate of its position, measured along the:

e:

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- a. observer's meridian.
- b. celestial equator.
- c. observer's horizon.
- d. ecliptic.

ANSWER:

b

- 67. In the right ascension coordinate direction, 1 hour corresponds to what equivalent angle?
  - a. 15°
  - b. 360°
  - c. 1°
  - d. variable depending on the time of year

ANSWER:

a

- 68. A comet moving northward from the equator across the sky can be described as having its:
  - a. declination increase with time.
  - b. right ascension decrease with time.
  - c. declination decrease with time.
  - d. right ascension increase with time.

ANSWER:

a

- 69. What is the angle, measured in degrees between two stars of RA =  $4^h$ , declination =  $0^\circ$ , and RA =  $8^h$ , declination =  $0^\circ$ ?
  - a. 4°
  - b. 0°, since declinations are equal
  - c. 120°
  - d. 60°

ANSWER:

d

- 70. The summer solstice position coordinates are RA =  $6^h 0^m 0^s$ , declination =  $+23^{\circ} 27'$ . What are the coordinates of the winter solstice?
  - a.  $RA = 18^h 0^m 0^s$ , declination =  $+23^\circ 27'$
  - b.  $RA = 0^h 0^m 0^s$ , declination =  $0^\circ 0'$ , by definition
  - c.  $RA = 12^h 0^m 0^s$ , declination =  $-23^\circ 27'$
  - d.  $RA = 18^h 0^m 0^s$ , declination =  $-23^\circ 27'$

ANSWER:

d

- 71. A star with  $RA = 4^h$  is in your meridian at a certain time. Which of the following stars will be on your meridian 2.5 hours from now?
  - a.  $RA = 6^h 30^m$

e:

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- b.  $RA = 4^h 2^m 30^s$
- c.  $RA = 1^h 30^m$
- d.  $RA = 4^h$ , since RA of a star does not change with time

ANSWER:

a

- 72. At what approximate value of declination was the Sun on March 21 this year?
  - a. 0°
  - b. 23.5°
  - c. 180°
  - d. no unique value

ANSWER:

a

- 73. At what approximate value of declination was the Sun on June 21 this year?
  - a.  $-23.5^{\circ}$
  - b. 90°
  - c. 0°
  - d. 23.5°

ANSWER:

d

- 74. At what approximate value of declination was the Sun on September 21 this year?
  - a.
- 23.5°
- b.
- 90° 180°
- c. d.
- 0°

ANSWER:

d

- 75. At what approximate value of declination was the Sun on December 21 this year?
  - a.
- 23.5°
- b.
- 90°
- c.
- $0^{\circ}$
- d.
- $-23.5^{\circ}$

ANSWER:

d

- 76. At what approximate value of right ascension was the Sun this year on March 21?
  - a. 1 hour
  - b. no particular value
  - c. 12 hours
  - d. 0 hours

ANSWER:

d

78. At what approximate value of right ascension was the Sun this year on September 21?  a. 18 hours b. 0 hours c. 6 hours d. 12 hours  ANSWER:  a. 0 hours b. 12 hours c. 18 hours d. 6 hours c. 18 hours d. 6 hours d. 6 hours  ANSWER:  c. 18 hours d. 6 hours  ANSWER:  c. 18 hours d. 6 hours  ANSWER:  c. 6 hours  ANSWER:  c. 6 hours  d. 12 hours c. 6 hours  b. 12 hours c. 6 hours c. 6 hours d. 12 hours d. 12 hours c. 6 hours d. 12 hours b. 13 hours d. 16 hours ANSWER:  c. 6 hours d. 10 hours b. 12 hours d. 12 hours d. 12 hours d. 12 hours c. 18 hours d. 12 hours d. 12 hours d. 12 hours d. 13 hours d. 6 hours	Name :			Class :	Dat e:
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d. 6 hours  ANSWER: c  80. The change in the right ascension of the Sun between June 21 and September 21 is approximately:  a. 0 hours.  b. 18 hours.  c. 6 hours.  d. 12 hours.  ANSWER: c  81. The change in the right ascension of the Sun between June 21 and December 21 is approximately:  a. 0 hours 0 minutes.  b. 12 hours.  c. 18 hours.  d. 6 hours.  ANSWER: b  82. The change in the declination of the Sun between December 21 and June 21 is approximately:  a. 90°.  b. 180°.  c. 47°.			b.	12 hours	
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d. 12 hours.  ANSWER:  c  81. The change in the right ascension of the Sun between June 21 and December 21 is approximately:  a. 0 hours 0 minutes.  b. 12 hours.  c. 18 hours.  d. 6 hours.  ANSWER:  b  82. The change in the declination of the Sun between December 21 and June 21 is approximately:  a. 90°.  b. 180°.  c. 47°.					
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d. 6 hours.  ANSWER:  b.  82. The change in the declination of the Sun between December 21 and June 21 is approximately:  a. 90°.  b. 180°.  c. 47°.					
ANSWER:  82. The change in the declination of the Sun between December 21 and June 21 is approximately:  a. 90°.  b. 180°.  c. 47°.					
<ul> <li>a. 90°.</li> <li>b. 180°.</li> <li>c. 47°.</li> </ul>	ANSWER:		•		b
<ul> <li>a. 90°.</li> <li>b. 180°.</li> <li>c. 47°.</li> </ul>	82. The change	e in the	declination of	f the Sun between December 21 and June 21 is approxi	mately:
b. 180°. c. 47°.	2. The change	- 111 1110			
c. 47°.					
(1 / 1 )			d.	23.5°.	

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ANSWER:			c
83. The change	in the declinat	ion of the Sun between March 21 and June	e 21 is approximately:
	a.	23.5°.	
	b.	90°.	
	c.	180°.	
	d.	47°.	
ANSWER:			a
a. b. c.		on angles between the north and south celending on the season.	estial poles is:
ANSWER:	<i>7</i> 0 .		c
	m are right as stars, ß Tauri west of south of north of	f the star β Tauri are right ascension 5 <sup>h</sup> 25 cension 7 <sup>h</sup> 44 <sup>m</sup> , declination +28°. Measure is: 3 Geminorum. β Geminorum. β Geminorum. Geminorum.	
ANSWER:			a
coordinates of th	en the two sta a. not b. eas c. sou	f the star $\gamma$ Geminorum are right ascension are right ascension $4^h$ $35^m$ , declination $+1$ rs, $\gamma$ Geminorum is: th of $\alpha$ Tauri. It of $\alpha$ Tauri.	
ANSWER:	d. we	st of α Tauri.	ь
		f the star γ Persei are right ascension 3 <sup>h</sup> 0.	

87. The celestial coordinates of the star  $\gamma$  Persei are right ascension 3<sup>th</sup> 03<sup>th</sup>, declination +53°; the coordinates of the star  $\rho$  Persei are right ascension 3<sup>th</sup> 03<sup>th</sup>, declination +39°. Measured along the shortest angle in the sky between the two stars,  $\gamma$  Persei is:

a. west of ρ Persei.

b. north of ρ Persei.

c. south of ρ Persei.

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	d. east of ρ Perse	i.	
ANSWER:	•		b
of the star $\lambda$ Ursa		nsion 10 <sup>h</sup> 16 <sup>m</sup> , declination +43°.	6 <sup>m</sup> , declination +24°; the coordinates Measured along the shortest angle in
b.	north of λ Ursae Maj		
c.	east of λ Ursae Majo		
d.	south of λ Ursae Ma	joris.	
ANSWER:			d
+31° 56N) and F  a. F  b. F  c. F  d. F  ANSWER:	Pollux (right ascension 7 <sup>h</sup> Pollux is slightly southeas Pollux is slightly northwere Pollux is slightly northeas Pollux is slightly southwere Pollux is slightly	44 <sup>m</sup> , declination +28° 5N). From the of Castor. St. of Castor. It of Castor.	(right ascension 7 <sup>h</sup> 33 <sup>m</sup> , declination m these coordinates you can see that:  a a attion +16° 28N) from a location in the
	here. Where is the star Bobelow Aldebaran and to	etelgeuse (right ascension 5 <sup>h</sup> 54 <sup>n</sup>	, declination +7° 24N)?
a. h	above Aldebaran and to		
c.	below Aldebaran and to		
d.	above Aldebaran and to	· ·	
ANSWER:		in right	a
91. The declinat	ion of Polaris, the North a. b. c. d.	Star, when viewed from a location 0°. 35°. 55°. 90°.	on 35° north of the equator, is:
ANSWER:	u.	,	d
92. Where must	you stand on Earth to ha	ve the celestial equator along you	ur horizon?

b.

c.

the South Pole

either pole

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	d.	the equator	
ANSWER:		•	c
93. Where mus	st you sta	nd on Earth to have the celestial equator pass	through your zenith?
	a.	the North Pole	
	b.	the South Pole	
	c.	either pole	
	d.	the equator	
ANSWER:			d
94. The appare	nt path o	f the Sun across the sky, day by day, through	out the year, is known as the:
	a.	celestial meridian.	
	b.	zenith.	
	c.	ecliptic.	
	d.	celestial equator.	
ANSWER:			c
95. The ecliptic	c is defin	ed as the:	
a. band o	of constel	lations through which the Sun and the Moon	move in the sky.
b. line in	the sky	hat is perpendicular to Earth's spin axis.	
c. line tra	aced in th	e sky by the Moon each month against the ba	ackground stars.
d. line tra	aced in th	e sky by the Sun over 1 year against the back	ground stars.
ANSWER:			d
96. The ecliptic	c crosses	the celestial equator at:	
a. tv	wo points	, known as solstices.	
b. or	ne point	only, known as the vernal equinox.	
c. th	ne meridi	an.	
d. tv	wo points	, known as equinoxes.	
ANSWER:			d
		background stars in daylight, how would the our motion on an orbiting Earth?	Sun appear to move against this
a.	1° per	day, from west to east	
b.	15° p	er hour, from east to west	
c.	1° per	day, from east to west	
d.	15° p	er hour, from west to east	
ANSWER:			a

98. If the daytime sky were not bright, in which direction would we see the Sun move along the ecliptic over the

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course of 1 year, relative to the background stars?

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- a. toward the east
  - b. toward the southwest
  - c. toward the west
  - d. toward the northwest

ANSWER: a

- 99. Which way are you moving with respect to the background stars because of the revolution of Earth *in its orbit around the Sun*?
  - a. westward
  - b. northeastward
  - c. northwestward
  - d. eastward

ANSWER:

- 100. At what average speed does the Sun appear to move across the sky with respect to the stars in order to move through one full circle in 1 year?
  - a. 15° per hour
  - b. about 1° per day
  - c. The Sun never appears to move with respect to the stars in the sky.
  - d. about 13° per day

ANSWER: b

- 101. What is the primary cause of Earth's seasons?
  - a. The orbit of Earth is an ellipse, so Earth is not always the same distance from the Sun.
  - b. Earth's rotation axis tilts with respect to the plane of its orbit around the Sun.
  - c. Earth's precession axis precesses (wobbles).
  - d. In accord with Kepler's second law, Earth moves faster during parts of its orbit around the Sun and more slowly during other parts.

ANSWER: b

- 102. The reason Earth experiences seasons is that:
  - a. Earth's rotation axis is not perpendicular to the ecliptic.
  - b. Earth's rotation axis is not perpendicular to the ecliptic, and the direction in which this axis points changes with time.
  - c. Earth is closer to the Sun during part of the year.
  - d. the Moon pulls on Earth from a distance that varies over the year.

ANSWER: a

- 103. If Earth's spin axis were perpendicular to Earth's orbital plane (the ecliptic plane), then the seasons and seasonal variation would be:
  - a. very little different from the present seasons.

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b. much faster (sho	orter seasons), but less severe.	
c. much more seven	re.	
d. nonexistent.		
ANSWER:		d
104. The tilt of Earth's spin ax	xis to the direction perpendicular to the ecliptic	e plane (known as the ecliptic pole)
	° and 23.5° over each 3-month season.	
b. 0°.		
c. 23.5°.		
d. 90°.		
ANSWER:		c
105. Earth would not have sea	asons if	
	vere perpendicular to its equatorial plane.	
	were perpendicular to its orbital plane.	
	vere perpendicular to its orbital plane.	
	were perpendicular to Earth's orbital plane.	
ANSWER:	LL L	c
106. Over a period of 6 month	ns, the tilt of Earth's spin axis with respect to the	ne background stars changes by an
a.	47°.	
b.		
c.	180°.	
d.	0°.	
ANSWER:		d
<ul><li>a. Earth is closest to the</li><li>b. sunlight falls more dir</li><li>c. Earth's equator is para</li><li>d. sunlight falls less dire</li></ul>	ere, summertime occurs when: Sun in its elliptical orbit. The rectly on this hemisphere, heating it more than allel to the plane of its orbit. The ctly on this hemisphere, spreading the heat out	t over a greater area.
ANSWER:		b

b.

c.

d.

108. Summertime in the northern hemisphere is when:
a. the Sun is closest to Earth.

the Moon is closest to Earth.

the Sun is closest to the ecliptic.

more direct sunlight shines on this hemisphere.

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ANSWER:		d
109. Winter in the northern hemispher a. Earth is farthest from the Sun b. sunlight falls most obliquely o c. Earth's axis is at its largest ang d. Earth is farthest from the eclip <i>ANSWER</i> :	in the elliptical orbit. n this region of Earth. gle with respect to the ecliptic plane be	ecause of precession.
110. The lowest amount of solar energy hemisphere on or about:  a. March 21, the end of with the begin discounty.  b. February 5, midwinter.  c. September 21, the begin discounty.  ANSWER:	inter. nning of autumn.	surface of Earth in the northern
111. At what time of the year will the hemisphere be the shortest?  a. noon, August 5, midsumr  b. noon, December 21, at th  c. dawn, June 21, at the beg  d. noon, June 21, at the begin	mer e beginning of winter inning of summer	at any site in the northern
a. midwinter, or early Jan b. midsummer, about Aug c. the first day of summer d. the first day of spring, a	uary gust 5 c, about June 21	c
during summer.  c. Earth is closer to the Sun in sund. the Sun is higher in the sky and	the days are shorter. sunlight passes through more atmosplemmer.	, -
ANSWER:		d

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114. Seasonal variations on a planet's surface occur because:

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· ·		e:

- a. the planet's distance from the Sun varies periodically over the orbital path.
- b. clouds alternately form and decay away in a periodic way.
- c. the planet's axis of spin is tilted with respect to the perpendicular to the orbital plane.
- d. volcanoes periodically cloud out the atmospheres of planets because of tidal interactions and distortions.

ANSWER:

- 115. When the northern hemisphere is experiencing winter:
  - a. Earth is closer to the Sun than it is during northern summer.
  - b. Earth is farther from the Sun than it is during the northern summer.
  - c. Earth is the same distance from the Sun as it is the rest of the year because Earth's orbit is circular with the Sun in the center.
  - d. the southern hemisphere is experiencing winter also.

ANSWER:

116. What does it mean to be "in the tropics," that is, between the Tropic of Cancer and the Tropic of Capricorn?

- a. The Sun is directly overhead at noon every day.
- b. The Sun is directly overhead on at least 1 day of the year.
- c. The stars do not appear to rise and set.
- d. Days and nights are of equal length all year long.

ANSWER: b

- 117. One required condition for seasons to occur is that a planet's:
  - a. spin axis be tilted with respect to the perpendicular to its orbital plane.
  - b. axis be perpendicular to its orbital plane.
  - c. atmosphere be thick.
  - d. distance from the Sun vary.

ANSWER:

- 118. The vernal equinox is one time of the year when the Sun:
  - a. crosses the Moon's orbital path in the sky.
  - b. crosses the celestial equator.
  - c. crosses the ecliptic plane.
  - d. is at its lowest point in the sky at midday.

ANSWER: b

- 119. The vernal equinox is that time of the year when:
  - a. the Sun crosses the equatorial plane, or celestial equator, moving north.
  - b. the Sun crosses the equatorial plane, or celestial equator, moving south.

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c. the Sun crosses the eclipt	ic plane.	
<u>-</u>	nt to the Sun in its elliptical orbit.	
ANSWER:	•	a
120. The equinoxes are located at	the intersections of the:	
a. ecliptic and the h		
b. ecliptic and the c	elestial equator.	
c. ecliptic and the N	Ioon's orbit.	
d. horizon and the c	elestial equator.	
ANSWER:		b
121. When the Sun is at one of the	e equinoxes:	
	length only for people on the equator.	
• •	night in one hemisphere of Earth and shorter	in the other hemisphere.
c. people on the equator have	-	
d. day and night are of equal		
ANSWER:		d
122. If you observe from a location morning throughout the year you wa. the Sun always rises in the		he horizon as it rises each
b. the Sun rises due east only	-	
-	progressively northward from December thro	ough June.
<u> </u>	progressively southward from December thro	_
ANSWER:	, ,	c
123. The Sun rises due east in the		
•	st day of spring and the first day of fall.	
b. on the equator on every		
	midsummer and midwinter.	
	of summer and the first day of winter.	
ANSWER:		a
as equinoxes. These points are the	d night are equal in length, the Sun is at one of intersections of which two planes in the sky	<u> </u>
a. ecliptic and celestia		
b. ecliptic and arctic of		
	and celestial equator	
d. celestial equator an	d ecliptic	

ANSWER:

d

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125. The autumnal equinox is that time of the year when:	
a. the Sun crosses the ecliptic plane, moving north.	
b. Earth is at the closest point to the Sun in its elliptical orbit.	
c. the Sun crosses the equatorial plane, moving south.	
d. the Sun passes through the galactic plane.	
ANSWER:	c
126. On the day of the vernal equinox (approximately March 21 each year), which of tholds?	he following conditions
a. The Sun rises at its most northerly point on the horizon on this day.	
b. Both day and night are almost exactly 12 hours long at all locations on Earth.	
c. Daylight is longest on this day.	
d. The Sun passes through an observer's zenith only on this day each year.	
ANSWER:	b
127. The approximate date March 21 represents the beginning of which season to peop Zealand?	ble living in New
a. winter	
b. summer	
c. autumn	
d. spring	
ANSWER:	С
128. The time of autumnal equinox, about September 21, is what season for Australian hemisphere?	ns in the southern
a. beginning of spring	
b. middle of winter	
c. middle of summer	
d. beginning of autumn or fall	
ANSWER:	a
129. At the summer solstice in the northern hemisphere, the Sun:	
a. is nearest to Earth.	
b. reaches its highest angle above the southern horizon for the whole year.	
c. is on the celestial equator.	
d. is at its lowest angle above the southern horizon at midday for the whole year.	
ANSWER:	b

130. As a result of the tilt of the spin axis of Earth to the plane of Earth's orbit (the ecliptic plane), sunrise in the

winter months in the mid-latitude northern hemisphere occurs in which direction in the observer's sky?

The Sun always rises due west.

a.

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b. s	southwest			
c. r	northeast			
d. s	southeast			
ANSWER:			d	
ever rise in the sou		<u> </u>	n parts by the east—west line, can the Sutude site in the northern hemisphere?	ın
b. No, since	the site is in the ne	orthern hemisphere.		
c. Yes, but o	nly for a few days	around midsummer.		
d. Yes, for m	ost of the year, si	nce the observing site is in the no	orthern hemisphere.	
ANSWER:			a	
a. Yes, for exb. No, since to c. Yes, but o	thern part of the s xactly half a year. the site is in the so nly for a few days	1	n parts by the east—west line, can the Sutude site in the southern hemisphere?	ın
ANSWER:		200 110 00001 / 111 g 2100 10 111 1110 200	a	
a. Yes, for exb. No, since to c. Yes, but o	thern part of the s xactly half a year. the site is in the so nly for a few days	<u> </u>	n parts by the east—west line, can the Sutude site in the southern hemisphere?  outhern hemisphere.	ın
134. In winter in th	ne southern hemis	phere, the Sun will rise on the:		
a.	northwestern	•		
b.	northeastern l	norizon.		
c.	southwestern	horizon.		
d.	southeastern l	norizon.		
ANSWER:			b	
135. The lowest lan	titude above whic	h, for at least 1 day per year, one	can see the Sun for a full 24 hours is	
-	a.	23.5°.		

b.

c.

52°.

66.5°.

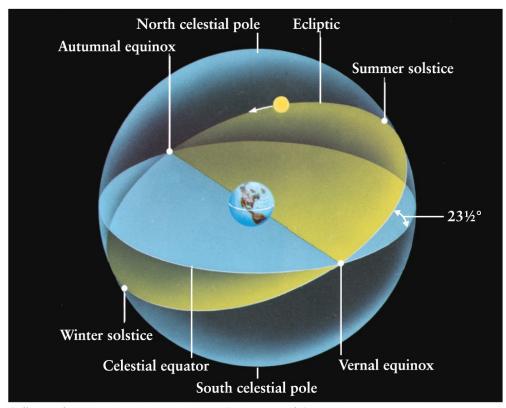
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		d.	90°.	
ANSWER:				c
following? a. 24 hours	s of sunligh	ıt	c Circle around the time of sumn	ner solstice will enjoy which of the
c. a period faint obj	of several ects in the		hich the Moon does not appear, a	allowing uninterrupted views of
ANSWER:	s of continu	ous darkness		a
horizon for a 24		od for at least a p	orthern or southern hemispheres to part of the year?	for the Sun to remain below the
b. nowher	e, since the	Sun is always v	visible at some time of the day an	nywhere on Earth
c. only at	90°, or at th	ne poles		
d. above a	bout 66.5°	latitude		
ANSWER:				d
138. Approxima spring at the No	•	ong will the Sun	remain above the horizon once	it first appears at the beginning of
	a.	18 hours		
	b.	6 months		
	c.	about 1 ho	ur	
	d.	12 hours		
ANSWER:				b
139. The Arctic	Circle is at	t a latitude of:		
a.	66.5°N	•		
b.	66.5°S.			
c.	23.5°N			
d.	a varial	ble average of 6	6.5°.	
ANSWER:		_		a
140 The Arctic	Circle is d	efined as a line	on Farth marking the southern he	oundary of the region where the Sun:
			or below the horizon.	Junuary of the region where the 5un.
	-	any time of the		
		winter and sumi		
	•		least 1 day of the year.	
ANSWER:	oc scen ioi	27 Hours on at	east I day of the year.	d
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- 141. The "Land of the Midnight Sun" is so named because:
  - a. the Sun is above the horizon for a full 24 hours at a certain time of the year.
  - b. the Sun passes overhead in this region at least once during the year.
  - c. twilight is bright and lasts all night through the summer months since the Sun never gets far below the horizon from these locations.
  - d. the full moon is always up whenever the Sun sets, maintaining light skies throughout the summer months.

ANSWER:

142. If you were standing on the equator, which of the following positions in the sky would pass directly over your head (i.e., through your zenith) at some time in one 24-hour period? (See Figure 2-15 of *Universe*, 11th ed.)



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- a. the position of the Sun at summer solstice
- b. the ecliptic pole, or perpendicular to the direction of the ecliptic plane
- c. the vernal equinox, or the zero point of the right ascension on the celestial equator
- d. the north celestial pole, or perpendicular to the direction of the celestial equator

ANSWER:

143. In what region of Earth would you have to be to have the Sun pass through your zenith at some time during Copyright Macmillan Learning. Powered by Cognero.

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the year?		
a. within the Arctic Circle		
b. at any latitude		
c. within $\pm -23.5^{\circ}$ of the equa	tor—the tropics	
d. only on the equator, nowher	e else	
ANSWER:		c
144. If the Sun passes through your zenith a. anywhere within 23.5° of	_ ,	you must be:
b. anywhere within 66.5° of	the equator.	
c. exactly on the equator.		
d. anywhere on Earth (no lim	nitation).	
ANSWER:		a
145. If you stand at latitude 10°N, how m zenith?	any times during the year will the	e Sun pass precisely through your
a. twice		
b. once		
c. never		
d. every day for a half a	year	
ANSWER:		a
146. Where would you have to be to see to a. at the North Pole	he south celestial pole on your ho	orizon?
b. about 1° away from the South 1	Pole, to allow for precession	
c. at the South Pole		
d. on the equator		
ANSWER:		d
147. During one complete year, an observ	er on the equator would be able t	o see what fraction of the overall
a. 50%		
b. a variable amount, depending	on the person's longitude	
c. a variable amount, depending	on which year	
d. 100%		
ANSWER:		d
148. During one complete year, an observ sky?	er at the South Pole would be abl	le to see what fraction of the overall

b. a variable amount, depending on which year

a. 50%

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c. 100 d. a va <i>ANSWER</i> :	% nriable amount, depending on	n the person's longitude	a
149. From the 1	North Pole:		
a. only	stars within 66.5° of the nor	th celestial pole can be seen.	
b. only	half the celestial sphere can	be seen on every clear night.	
c. only	stars 23.5° above the celestic	al equator can be seen.	
d. the w	hole celestial sphere is visib	le at some time during the year.	
ANSWER:			ь
<ul><li>a. abou</li><li>b. North</li><li>c. South</li></ul>	t 1° away from the South Po h Pole h Pole	to see the north celestial pole in sle, to account for precession	your zenith?
d. equa	tor		
ANSWER:			b
-	re standing on the South Pole would you see the Sun all d	<u> </u>	your zenith at the time of the vernal
a.	1 .		
b.	23.5° above the horizon	on	
c.	well below your horiz	on	
d.	at your zenith		
ANSWER:			a
your horizon (a a. 90° b. 0°; it w	at midday, of course)?		et angle the Sun would reach above vays in darkness.
a. 23.3 ANSWER:			d
11,077 1210,			u
153. If you wer Sun to be at mi	dday?	e at the time of the autumnal equ	uinox, where would you expect the
a.	on your horizon		
b.	well below your horize	on	

c.

d.

at your zenith

23.5° above your horizon

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	· ·	e:

ANSWER:

154. Where would you expect to see the Sun in your sky if you were at the North Pole at the beginning of fall (about September 21)?

- a. at your zenith
- b. about 23.5° above your horizon all day
- c. on the horizon
- d. below your horizon all day

ANSWER:

c

- 155. What would be the position and motion of the Sun on December 21 from the South Pole on Antarctica?
  - a. It would rise in the east at 6 A.M. and set in the west at 6 P.M., reaching 47° above the horizon at midday.
  - b. It would pass across the sky from the horizon at midnight to reach an angle of 23.5° above the horizon at midday and then return to the horizon.
  - c. It would remain below the horizon for the whole 24 hours.
  - d. It would move completely around the sky in 24 hours while maintaining an angle of 23.5° above the horizon.

ANSWER:

156. You view the night sky for a period of 1 hour. What changes during this hour?

- a. the right ascension of a particular star, like Castor in Gemini
- b. the declination of a particular star, like Pollux in Gemini
- c. the right ascension of the star at your zenith
- d. the position of Polaris

ANSWER: c

- 157. Suppose Earth's rotation axis were straight up and down with respect to the plane of the ecliptic and not tilted. How would a person exactly at the north pole view the Sun?
  - a. The Sun would always be below the horizon and never visible.
  - b. The Sun would always be visible on the horizon.
  - c. The Sun would be visible for half the year, as is the case now.
  - d. The Sun would be visible for more than half the year, but not for the whole year.

ANSWER: b

- 158. Suppose Earth's rotation axis were straight up and down with respect to the plane of the ecliptic and not tilted. Each of the following statements would be true except one. Which one would *not* be true?
  - a. The stars would rise along paths perpendicular to the horizon (rather than slanted) when viewed from middle latitudes.
  - b. Seasons, as we presently experience them, would not exist.
  - c. Every point on Earth would experience 12 hours of sunlight and 12 hours of darkness each day.

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d. In general, the declinations of standard ANSWER:	ars would be different from their pres	sent values.
159. What is the definition of the vernal a. March 21	•	. 1. 01
c. the date when the Sun first enter		
d. the point on the ecliptic where the ANSWER:	he Sun crosses the celestial equator f	from south to north
<ul><li>160. What is the right ascension of the S</li><li>a. 0 hours</li><li>b. 12 hours</li></ul>	Sun?	
	ery slowly because of precession cantly over a day	d
161. What is the declination of the Sun?  a. 0°  b. 90°	)	
<ul><li>c. a number that changes only ve</li><li>d. a number that changes significant</li></ul>	ery slowly because of precession	
ANSWER:	cultify over a day	d
<ul><li>162. What is the declination of Polaris?</li><li>a. 0°</li><li>b. 90°</li></ul>		
<ul><li>c. a number that depends on the l</li><li>d. a number that changes signific</li></ul>	latitude from which you are observin antly over a day	g
ANSWER:	, and the second	b
<ul><li>163. In its motion across the sky against move about:</li><li>a. 0.5° per day, its own diame</li></ul>	_	of a month, the Moon appears to
b. 1.0° per day, twice its diam	eter, from west to east.	
<ul> <li>c. 0.5° per hour, its own diame</li> <li>d. 0.5° per hour, its own diame</li> </ul>		
ANSWER:		d

164. The zodiac is a:

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	•	Δ'

- b. constellation representing a boat in the sky.
- c. band of sky extending 8° on either side of the ecliptic.
- d. band of sky 8° wide centered on the ecliptic.

ANSWER: c

165. The total width of the zodiac surrounding the ecliptic plane in the sky is approximately:

- a. 8°.
- b. 16°.
- c. 23.5°.
- d. 1°.

ANSWER: b

166. Precession is:

- a. the motion of Earth along its orbital path.
- b. a very slow conical motion of Earth's axis of rotation.
- c. the occasional reversal of the direction of spin of Earth.
- d. the daily spinning motion of Earth.

ANSWER: b

#### 167. Precession is:

- a. the slow coning motion of the spin axis of Earth, similar to that of a spinning top.
- b. the daily spinning motion of Earth, producing the apparent motion of the Sun and the stars.
- c. another name for a parade.
- d. the motion of Earth along its orbital path during a year.

ANSWER: a

- 168. What is the significance of Polaris, the North Star?
  - a. It has been chosen as the zero point for declination measurements because it will always be at the pole of Earth's axis, i.e., due north.
  - b. The direction from Earth to Polaris is used as the zero line of prime meridian for measurements of right ascension.
  - c. Polaris is near the center of the curve traced out by the rotation axis as a result of the precession of the rotation direction.
  - d. Polaris is almost on the curve traced out by the rotation axis as a result of the precession of the rotation direction.

ANSWER:

- 169. The direction of the axis around which Earth rotates:
  - a. always points toward Polaris, the North Star.
  - b. precesses (wobbles) and takes about a century to go around once.
  - c. precesses and takes many thousands of years to go around once.

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		e:

d. always tilts slightly toward the direction of the Sun.

ANSWER:

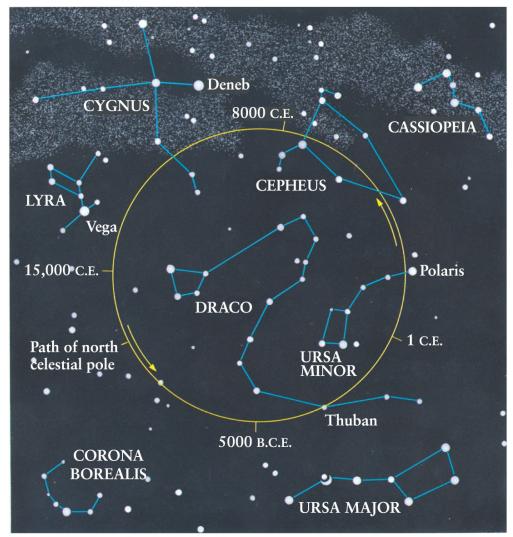
- 170. Precession of Earth's spin axis results in a:
  - a. cyclic variation over a period of 1 year of the constellations visible at night from Earth.
  - b. daily shift in the position of the observer's zenith relative to the celestial equator (for an observer at a fixed location on Earth).
  - c. gradual shift in the angle between the ecliptic and the celestial equator.
  - d. gradual shift of the vernal equinox along the ecliptic.

ANSWER:

- 171. A science fiction writer, writing a story about inhabitants on Earth in 14,000 C.E. who have survived a disaster that included the loss of modern navigational aids, describes them traveling due north across barren wastes by walking toward Polaris, the Pole Star. What is wrong with this situation?
  - a. Polaris will have moved away from due north since it is moving rapidly with respect to surrounding stars.
  - b. Polaris will no longer be due north, because of Earth's precession.
  - c. By that time, Polaris will be due south, not due north, because of the reversal of Earth's spin axis.
  - d. Polaris will no longer be visible since its lifetime is only a few thousand years.

ANSWER: b

172. Polaris is our "pole star" at the present time. At approximately what time in history or in the future would the star Thuban be our "pole star?" (See Figure 2-20 of *Universe*, 11th ed.)



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- a. never, since Polaris will always be our "pole star" when viewed from Earth
- b. 1 C.E.
- c. 3000 C.E.
- d. 3000 B.C.E.

ANSWER:

173. The position in Earth's orbit around the Sun, which presently corresponds to winter solstice, will instead correspond to summer solstice in about:

- a. 6500 years.
- b. 13,000 years.
- c. 19,500 years.
- d. 26,000 years.

ANSWER: b

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	· ·	e:

- 174. Precession of Earth's axis of rotation is caused by:
  - a. changes in the rate of rotation (length of the day) of Earth caused primarily by the gravitational pull of the Moon.
  - b. changes in the shape of Earth's orbit due to the gravitational pull of the Moon.
  - c. the gravitational pull of the Moon and the Sun on the equatorial bulge of Earth.
  - d. changes in the shape of Earth's orbit due to the gravitational pull of the Sun.

ANSWER:

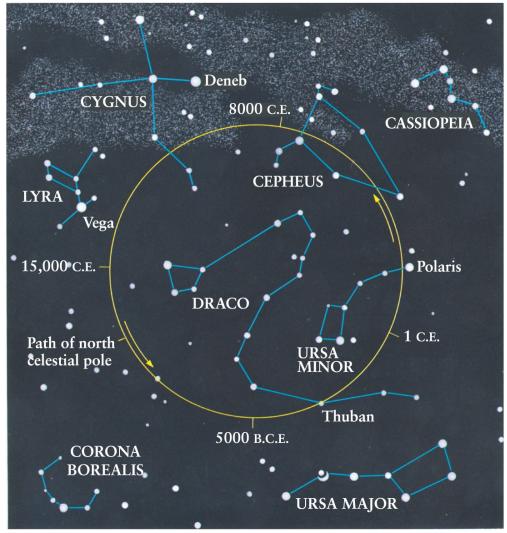
- 175. The reason for the slow movement of the vernal equinox through the sky against the background stars over long periods of time is the:
  - a. precession of the spin axis of Earth.
  - b. overall movement of local stars in the sky.
  - c. movement of the Sun in the Milky Way Galaxy.
  - d. motion of Earth in its orbit.

ANSWER: a

- 176. The phenomenon of precession of Earth's spin axis is caused by the:
  - a. varying intensity of sunlight on Earth throughout the year.
  - b. tidal ebb and flow of ocean waters on Earth.
  - c. variation of the spin rate of Earth.
  - d. gravitational pull of the Moon and the Sun on Earth's equatorial bulge.

ANSWER:

177. If the polar axis of Earth precesses through a full circle (See Figure 2-20 of *Universe*, 11th ed.) in 26,000 years, how many years will it take for the line between the axis and the center of the circle to move through 1°?



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a. 720

b. 0.014

c. 7.2

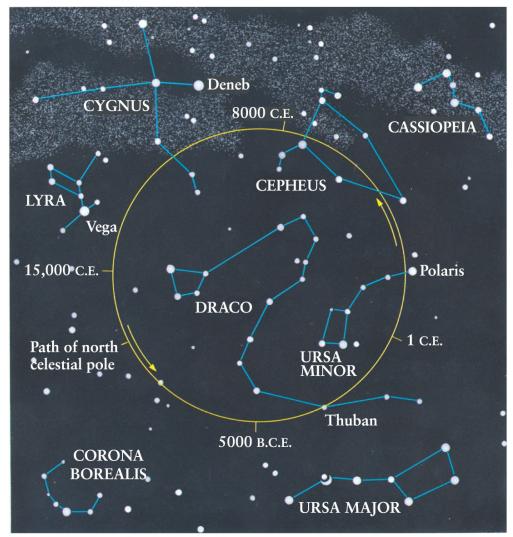
d. 72

ANSWER:

178. To what constellation will the north celestial pole be closest in the year 14,000 C.E.? (See Figure 2-20 of *Universe*, 11th ed.)

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- a. Draco
- b. Lyra

ANSWER:

- c. Cepheus
- d. Ursa Major, since the north celestial pole never moves

ANSWER: b

179. The precessional motion of the north celestial pole of Earth over a period of 26,000 years is a circle of 47° diameter in the northern sky. The equivalent motion of the south celestial pole is:

- a. a circle of 47° diameter, covered in 26,000 years.
- b. a much smaller circle, covered in 26,000 years.
- c. a circle of 47° diameter, covered in a much shorter time, about 1000 years.
- d. the south celestial pole does not move at all during this precession.

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180. The time the of precession is:	-	se before the spin as	xis of Earth points toward	the present pole star again as a result
a.	9 yea	rs.		
b.	at lea	st 1 million years.		
c.	13,00	0 years.		
d.	26,00	0 years.		
ANSWER:				d
181. The slow c circle in a period		rn of the spin axis of	f Earth's precession will m	nove the end of this axis in a complete
	a.	26 million years.		
	b. 2	26,000 years.		
	c. 2	2600 years.		
	d.	1 year.		
ANSWER:				b
182. The position	on of the vei	mal equinox in the s	ky is at present in the con	stellation:
	a.	Ursa Major.	, <sub>[</sub>	
	b.	Pisces.		
	c.	Aquarius.		
	d.	Aries.		
ANSWER:				b
183. Do the astr night?	onomical co	oordinates of right a	scension and declination of	of a star change systematically night by
a. They	do not chan	ge, since they are po	sitions on a fixed star cha	rt.
b. Yes, b	ecause of the	ne motion of Earth in	n its orbit around the Sun.	
c. Yes, b	ecause of the	ne bending of light b	y Earth's atmosphere.	
d. Yes, b	ecause of the	ne precession of Ear	th's spin axis.	
ANSWER:				d
to the rising of t of the equinoxes a. The align	he Sun at vos affect thes nments of s	ernal equinox or son e alignments in a bu uch a building shoul	ne other important astrono hilding, say, 2000 years old d still be perfect.	
b. The align	nments will	vary slightly, but al	tter a mere 2000 years the	discrepancy in alignment will be

ANSWER:

c. The alignments will vary significantly after 2000 years, but only for buildings within 23.5° of the

d. The alignments will vary significantly after 2000 years for buildings anywhere on Earth.

d

equator.

almost too small to measure.

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	rly the same as:  poles and the zenith (for an observer  poles and the vernal equinox (for an	*
ANSWER:		b
186. In 1 year the apparent path of the S  a. only the 12 zodiac constellations  b. the 12 zodiac constellations plus	s defined by the ancients.	tars passes through:
-	c constellations: The path no longer	passes through Aries.
ANSWER:		ь
a. the right ascension of a typical st.  b. the declination of a typical st.  c. the date of the vernal equinor	al star ar	Earth's spin axis?
d. the definition of the zero of of the		d
188. Which direction in an observer's sk  a. the zenith  b. the north celestial p  c. due south on the ho  d. the vernal equinox	pole	an?
ANSWER:		d
	and south celestial poles.  nd west positions on his or her horizoles and the vernal equinox position	
ANSWER:	•	a
190. A given star in the sky will reach it a. zodiac.  b. ecliptic plane		erver when it passes through the:

c.

meridian.

·	:e:	
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d.	celestial equator.	
ANSWER:	1	c
191. An observer's ce	elestial meridian is the:	
a. are joining the	he north and south celestial poles through the observer's zenith.	
b. extension of	f the horizon onto the sky.	
1	th's orbit extended onto the sky.	
	Earth's equator onto the sky.	
ANSWER:		a
192. The line on the s	sky joining the north and south celestial poles through a person's zenith is l	known as the
a.	meridian.	
b.	celestial equator.	
C.	zodiac.	
d.	horizon.	_
ANSWER:		a
mid-latitudes, roughl	ng Leo (the lion) extends for about 30° along and close to the celestial equally how long does it take Leo to rise above the horizon?  a. 5 hours  b. 30 seconds  c. 30 minutes  d. 2 hours	
ANSWER:		d
<ul><li>a. at a varying t</li><li>b. about 4 minu</li><li>c. about 4 minu</li></ul>	he sky closely night by night, you would note that a particular star rises: time every night, sometimes earlier, sometimes later than a specified time. Ites later every night.  Ites earlier every night.  Item every night.	0
ANSWEK.		c
<ul><li>a. about</li><li>b. about</li><li>c. about</li></ul>	t the Sun), when viewed from low and mid-latitudes, rises in the east: t 1 hour later each evening. t 4 minutes later each evening. t 4 minutes earlier each evening. e same time each evening.	
ANSWER:		c
196. The Big Dipper, solar time?	, or Ursa Major, will return to the same position in an observer's sky in wha	at time period in

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	a.	24 hours	s 4 minutes			
	b.	365.25	lays			
	c.	23 hours	s 56 minutes			
	d.	24 hours	s exactly			
ANSWER:			·		c	
		contact wit		ter. On the equator, appro the Sun completely belo	ximately how long does it take for th w that horizon?	ne
			2 minutes			
	c. d.		about 1 hour			
ANSWER:	u.		about I noui		c	
to Earth in	its orbit. If see from Ve 0.25 seco about 4 s about 4 r	You were enus first reonds seconds minutes	watching Venus	s set in the west over a clo zon to Venus completely	in (1/60°) when it is at the closest point ear horizon (e.g., ocean), how long setting below the horizon?	IIIt
ANSWER:					b	
night, what a. 10: b. Th c. 10:	time would :04 P.M. is star will :00 P.M.	ld the same	e person (at the		10 P.M. (2200 hours) on a particular ar cross the horizon on the next night after 1 year.	?
	56 P.M.					
ANSWER:					d	
	ely what the		_	Procyon will just be rising r rising on Christmas Day	g on the eastern horizon. At (24 days later)?	
		d.	8:24 P.M.			
ANSWER:					d	

201. At 10 P.M. on December 1, the bright star Procyon is seen to rise on the eastern horizon. At approximately

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what time will this star rise 7 days later, on December 8?

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- a. 10:28 P.M.
- b. 9:53 P.M.
- c. 9:32 P.M.
- d. 10:00 P.M.

ANSWER:

- 202. A sundial is not considered to be a good timekeeper because:
  - a. the sky is often cloudy.
  - b. Earth's rotation rate changes throughout the year.
  - c. the Sun's large angular diameter produces a fuzzy shadow.
  - d. Earth's orbital speed around the ecliptic is variable.

ANSWER:

- 203. Time, as indicated by an uncorrected sundial, can differ from true time by as much as 15 minutes at certain times of the year. This is because:
  - a. Earth moves non-uniformly along an orbit inclined to the celestial equator.
  - b. the Sun moves with respect to the background stars because of our motion on an orbiting Earth.
  - c. the time zones are poorly defined on Earth's surface, leading to time errors that vary through the year.
  - d. Earth's rotation rate varies throughout the year.

ANSWER:

- 204. Sidereal time is the more fundamental time, since it is a measure of the true rotation rate of Earth. Why then do we govern our lives by solar time rather than sidereal time?
  - a. We cannot divide the day into 24 equal hours of sidereal time.
  - b. In sidereal time, the Sun would reach the meridian early or late, sometimes by as much as 17 minutes at certain times of the year.
  - c. Different clocks tick at different rates depending on latitude in sidereal time.
  - d. We wish to remain in time with the Sun's illumination on Earth, with high Sun angle at about midday every day.

ANSWER:

- 205. The difference between 1 second of sidereal time and 1 second of solar time is:
  - a. very small but variable because of the variable motion of Earth in its orbit.
  - b. 0, since they are defined to be equal.
  - c. a very small but finite and fixed interval of time.
  - d. very large, because 1 sidereal second measures a fraction of 1 year while 1 solar second measures the same fraction of 1 solar day.

ANSWER:

206. At a particular time, the vernal equinox passes through the celestial meridian of an observer. What is the

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sidereal time at that	site?		
a. 12 hou	rs (noon) exactly		
b. 0 hours	s (midnight) exactly		
c. variabl	e, depending on the observ	ver's latitude	
d. variabl	e, depending on the time of	of year	
ANSWER:			b
207. The beginning passage of the:	of a sidereal day (at midni	ght, sidereal time) at any loca	ation on Earth is defined by the
a. pole star	through the lower meridia	n at that site.	
b. Sun thro	ugh the position of the ver	nal equinox.	
c. vernal ec	quinox through the upper r	neridian at that site.	
d. pole star	through the upper meridia	n at that site.	
ANSWER:			c
b. c. d. <i>ANSWER</i> :	sidereal time only. UTC only. any of these.		c
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a.	the pole star.	ere are minimized when you o	observe in the direction of:
ь. b.	the eastern horizon.		
c.	the upper meridian.		
d.	the lower meridian.		
ANSWER:			c
210. Sidereal clocks	are based on the position	of:	
a.	the Sun.		
b.	the vernal equinox.		
c.	Polaris.		
d.	the summer solstice.		
ANSWER:			ь
211. A calendar year	r, from vernal equinox to v	vernal equinox, is called a:	
а	tropical year.		
t	o. sidereal year.		

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c. solar year.

d. vernal year.

ANSWER:

- 212. Which of the following were leap years according to the calendar then in use: 1000, 1492, 1600, 1776?
  - a. just 1600
  - b. just 1492 and 1776
  - c. all except 1600
  - d. all of these years

ANSWER:

- 213. The direction of Earth's rotation about its axis is the same as the direction of its revolution about the Sun. How does a solar day compare to a sidereal day on Earth?
  - a. A solar day is always longer.
  - b. A sidereal day is always longer.
  - c. They are always the same length.
  - d. A sidereal day is longer when Earth is farther from the Sun (northern summer), but a solar day is longer when Earth is closer to the Sun (northern winter).

ANSWER: a

- 214. A solar day is the time it takes Earth to rotate around its axis between two consecutive solar positions (i.e., high noon to high noon or sunset to sunset). A sidereal day is the time it takes Earth to rotate around its axis between two consecutive positions of a distant star (i.e., Vega on the eastern horizon to Vega again on the eastern horizon). Which is longer?
  - a. A solar day is always longer.
  - b. A sidereal day is always longer.
  - c. They are always the same length.
  - d. A sidereal day is longer when Earth is farther from the Sun (northern summer), but a solar day is longer when Earth is closer to the Sun (northern winter).

ANSWER: a

- 215. The tropical year is about 20 minutes shorter than the sidereal year. What does this tell you about precession?
  - a. The vernal equinox moves eastward along the ecliptic.
  - b. The vernal equinox moves westward along the ecliptic.
  - c. The vernal equinox remains fixed on the ecliptic, but the Sun progressively changes its apparent position against the background stars from one sidereal year to the next.
  - d. This time difference has nothing to do with the precession of the equinoxes.

ANSWER: b